# **Thermal Conductivity and Thermal Diffusivity of Ferrosilite under High Temperature and High Pressure**

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ABSTRACT: **Orthopyroxene is an important constitutive mineral in the crust and the upper mantle. Its thermal properties play a key role in constructing the thermal structure of the crust and the upper mantle. In this study, we developed a new method to synthesize polycrystalline ferrosilite, one endmember of orthopyroxene, via the reaction of**  $FeO + SiO<sub>2</sub> \rightarrow FeSiO<sub>3</sub>$ **. We found that the** *P***-***T* **condition of 3 GPa and 1 273 K is suitable to synthesize dense ferrosilite samples with low porosity. We employed the transient plane-source method to investigate the thermal conductivity**  $\kappa$  and thermal diffusivity  $D$ of synthetic ferrosilite at 1 GPa and 293–873 K, of which,  $\kappa = 1.786 + 1.048 \times 10^{3} T^{-1} - 9.269 \times 10^{4} T^{-2}$  and  $D = 0.424 + 0.223 \times 10^{3} T^{1} + 1.64 \times 10^{4} T^{2}$ . Our results suggest phonon conduction should be the domi**nant mechanism at** *P***-***T* **conditions of interest since the thermal conductivity and the thermal diffusivity of ferrosilite both decrease with increasing temperature. The calculated heat capacity of ferrosilite at 1 GPa increases with temperature, which increases with increasing temperature with about 10% per 100 K (<500 K) and 4% per 100 K (>500 K). Iron content of an asteroid significantly influences its thermal evolution history and temperature distribution inside. It is expected that the mantle temperature of the Fe-rich asteroid will be higher and the Fe-rich asteroid's cooling history will be longer.**

KEY WORDS: **ferrosilite, high pressure, thermal conductivity, thermal diffusivity, synthesis.**

# **0 INTRODUCTION**

Pyroxene is considered to be one of the main constituent minerals in the upper mantle (Ringwood, 1975), which includes two subgroups of orthopyroxene and clinopyroxene. Ferrosilite (FeSiO<sub>3</sub>) is one end-member of orthopyroxene. Enstatite (Mg- $SiO<sub>3</sub>$ ) and ferrosilite can form ideal solid solution through the Mg-Fe substitution. Although the iron number  $[100 \times Fe/(Fe +$ Mg)] of the shallower upper mantle is about 10 (Ono and Oganov, 2005; Gaul et al., 2000), not only in the deeper mantle where composition heterogeneity has been observed (Ringwood, 1991), but also in terrestrial planes whose Fe concentration is distinguished from that of the Earth, e.g., the Mars (Khan et al., 2018). The Fe concentration of some asteroids surface was also found to range from 49 to 70 (Sanchez et al., 2014; Sunshine et al., 2007). The iron content of orthopyroxenes could be different. It has been known that iron content greatly influences the physical properties of orthopyroxene (Zhang and Yo‐ shino, 2016; Hofmeister, 2012; Stalder, 2004), and also influenc‐ es the thermal conductivity and thermal diffusivity of olivine (Zhang Y Y et al., 2019). In order to illuminate iron content ef‐ fect on the thermal conductivity and thermal diffusivity of or‐ thopyroxene, it is necessary to survey the thermal properties of ferrosilite, which has not been reported ever before.

To synthesize ferrosilite, high-pressure and hightemperature technique is generally required. There are mainly two methods to synthetize ferrosilite: One method employed mixture of  $SiO_2$ , Fe,  $O_3$ , Fe and 1 wt.% ultrapure water to synthesize ferrosilite between  $1\ 000 - 1\ 200\ ^\circ\text{C}$  at 3 GPa (Xu et al., 2020). The employment of water will inevitably inject hydrogen into the crystal structure of ferrosilite at high pressure, and may strongly affect the thermal properties of the minerals (Chang et al., 2017). Another method employed high-temperature gas flow furnace to synthesize fayalite firstly from mixture of  $Fe<sub>2</sub>O<sub>3</sub>$  and SiO<sub>2</sub>, and then employed high-pressure apparatus to synthesize ferrosilite finally from mixture of fayalite and SiO<sub>2</sub> at high temperatures (Kung and Li, 2014; Giuli et al., 2002; Akimoto et al., 1964; Lindsley et al., 1964). This method is rel‐ atively complicated and the stability field of ferrosilite is not well determined yet (Giuli et al., 2002; Lindsley et al., 1964; Bowen and Schairer, 1935). Therefore, it is necessary to devel‐ op the synthesis method to overcome the above disadvantages and define effects of pressure and temperature.

In this study, we directly synthesize well-sintered ferrosi‐ lite with low porosity using the mixture of FeO and  $SiO<sub>2</sub>$  under high pressure and high temperature. Via the high quality poly‐ crystals, we measured the thermal conductivity (*κ*) and thermal diffusivity (*D*) of ferrosilite at 1 GPa and 293–873 K. Based on

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the present experimental data, we calculated the temperature distribution and the thermal evolution of the Fe-rich and Fepoor asteroids.

## **1 EXPERIMENTAL METHOD**

## **1.1 Experimental Instruments**

A YJ-3000 multi-anvil apparatus, installed at the High-Temperature and High-Pressure Laboratory of the Institute of Geochemistry, Chinese Academy of Sciences, Guiyang, was used for synthetic experiments and thermal conductivity experiments. The pressure was calibrated from the phase transition of metallic bismuth (ambient temperature at 2.54 GPa) and halide melting experiment (Fu et al., 2019). The temperature is monitored by a K-type (NiCr-NiAl) thermocouple. The approximate errors of pressure and temperature estimation are 0.1 GPa and 0.5 K, respectively.

#### **1.2 Synthesis of Ferrosilite**

High-purity FeO powder with a grain size of about 10 μm together with SiO<sub>2</sub> powder (5  $\mu$ m or 30 nm) were used the starting materials. The SiO, powder was baked at 1 273 K for 1 h in an oven before be weighted. The FeO and SiO, powders were weighed in equal mole ratio and well mixed in an agate mortar with acetone for 2 h. Then the mixture was dried in a vacuum container for 48 h. Finally, the mixture was compressed to a cylindrical shape by a hand-press and loaded into the sample cap‐ sule. Three materials, graphite, iron and molybdenum, were used as the sample capsules to investigate the effect of capsule on the synthesis of ferrosilite. The entire mixing and assembling process was operated in a transparent box filled by argon to prevent the oxidization of FeO. Under suitable conditions, ferrosilite can be synthesized through reaction

$$
\text{FeO} + \text{SiO}_2 \rightarrow \text{FeSiO}_3 \tag{1}
$$

Ferrosilite aggregates for thermal conductivity measurements were sintered with a Mo capsule in a multi-anvil appara‐ tus. After recovery, disks with  $\sim$ 1.0-mm thickness and  $\sim$ 5.0mm diameter were cored and used for subsequent measurements. The detailed experimental conditions and results are summarized in Table 1.

#### **1.3 Ferrosilite Characterization**

The micro-texture of the recovered sample was observed by a JSM-7800F scanning electron microscopy (SEM). The fer‐ rosilite synthesized under the above conditions has been confirmed by the Raman technique (Fig. 1). The SEM observation shows that the sample contains nearly pure ferrosilite with a grain size of about 10 μm (Fig. 2a, YJ-123). The ferrosilite is dense and machinable with low porosity (Fig. 3), which is suitable for the measurements of thermal properties under high *P*-*T* conditions.

## **1.4 Thermophysical Properties Measurement**

The measurement of thermophysical properties adopts the transient plane-source method (Osako et al., 2004; Dzhavadov, 1975). The assembly of thermal conductivity experiment is shown in Fig. 2. The detailed description for the cell assembly and method is basically the same as Zhang B H et al. (2019).



Figure 1. Raman spectra of synthetic ferrosilite. A laser beam with a wavelength of 514 nm was used to generate the Raman spectra of the sample. Totally 5 points were collected at 100 μm interval. Reference data are from Stalder et al. (2009).

Three disk samples with the same dimensions are piled face to face vertically at the center of the sample capsule. A planar impulse heater made of nickel was placed on one interface between the disks, while the thermocouple junction was set on the opposite interface. The sample was insulated from the graphite heater by MgO. Two nickel disks directly contacting with the sample were used as the heat sinks to satisfy the constant-temperature boundary.

During cooking, each pulse heating will cause the temper‐ ature fluctuation (∆*T*) of the sample, which can be expressed as

$$
\Delta T = A \sum_{n=1}^{\infty} \frac{1}{n^2} \sin \frac{n\pi}{3} \sin \frac{n\pi x}{d} \cdot \exp\left(-n^2 B t\right) \left[\exp\left(n^2 B \tau\right) - 1\right] (t > \tau) \tag{2}
$$

where  $x$  is the distance between impulse heater and thermocouple (m), *d* is the overall thickness of the three samples (m),  $\tau$  is the pulse heating duration  $(s)$ , and  $t$  is the time from the onset of heating (s). The parameters *A* and *B* are defined as follows

$$
A = \frac{2Qd}{\pi^2 \kappa S}, B = \frac{\pi^2 D}{d^2} \tag{3}
$$

where  $Q$  is the power used for impulse heating (W),  $S$  is the area of impulse heater (m<sup>2</sup>),  $\kappa$  is the thermal conductivity (W·m<sup>-1</sup>· K<sup>-1</sup>), and *D* is the thermal diffusivity (mm<sup>2</sup>·s<sup>-1</sup>). Since Eq. (2) for Δ*T* can reach a rapid convergence, a summation of *n* up to 15 should yield sufficiently accurate values (Yoneda et al., 2009; Osako et al., 2004), whose reliability has been confirmed by a finite element simulation (Yoneda et al., 2009).

The effect of heterogeneous temperature on results is negligible since with 8–16 W impulse power, the temperature distur‐ bance across sample caused by impulse heating is only 3–6 K. Sample thickness and area during measurement was corrected according to the Birch-Murnaghan equation of state (EoS) (Hugh-Jones and Angel, 1997). Area change of impulse heater was corrected according to the equation  $S_{(P,T)}/S_0 = (V_{(P,T)}/V_0)^{2/3}$  used by Wang et al. (2014), where  $S_{(P,T)}$  and  $V_{(P,T)}$  are impulse heater ar-

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No.	T(K)	$P$ (GPa)	Duration (h)	Starting material	Capsule	Reaction products
Kawai-5	1 5 7 3	3	20	$SiO2 + FeO$	Mo	Quench $(Fs + FeO) + Qtz + Fe$ -Mo alloy
YJ-6	1 3 7 3	1.5	20	$SiO2 + FeO$	Mo	$FeO + Otz$
$YJ-7$	1 5 2 3	1.5	9	$SiO, + FeO$	Mo	Quench $(Fa) + Qtz + Fe$ -Mo alloy
$YJ-29$	1 573	3	20	$SiO2 + FeO$	Graphite	Quench $(Fs + FeO) + Otz$
$YJ-30$	1 5 2 3	3	20	$SiO2 + FeO$	Graphite	Quench $(Fs + FeO) + Otz$
$YJ-35$	1473	3	20	$SiO2 + FeO$	Graphite	Quench $(Fs + FeO) + Otz$
$YJ-53$	1 2 7 3	3	20	$SiO2 + FeO$	Fe	$Fa + Fs$
$YJ-55$	1 2 7 3	2	20	$SiO3 + FeO$	Graphite	Fs
$YJ-56$	1 2 7 3	3	20	$SiO3 + FeO$	Graphite	Fs
$YJ-58$	1 3 7 3	2	15	$SiO3 + FeO$	Graphite	$Fa + Qtz$
$YJ-96$	1 2 7 3	3	20	$SiO, + FeO$	Graphite	$Fs + FeO + Otz$
$YJ-170$	1 2 7 3		20	$SiO3 + FeO$	Graphite	Quench $(Fa) + Qtz$
$YJ-123$	1 273	3	$\overline{2}$	Fs powder	Mo	Fs aggregates

**Table 1** Experimental conditions and results for the synthesis of ferrosilite

\*. Nano-scale SiO, was used; Fa. fayalite; Fs. ferrosilite; Qtz. quartz.

ea, and sample volume at high-pressure high-temperature condi‐ tions, respectively; and  $S_0$ , and  $V_0$  are impulse heater area, and sample volume in ambient conditions, respectively. In the study, the total experimental systematic error is approximately 5% for both *D* and  $\kappa$ , which is mainly caused by errors associating with pulse heating power *Q*, sample thickness *d*, heating area *S*, and the least square fitting parameters of *A* and *B* in Eq. (3). Adopting the transient plane-source method, we measured *κ* and *D* of ferrosilite (YJ-123) at 1 GPa and 293–873 K (Table 2). In the experiment, the pressure was raised to the target val‐ ue at 1 GPa/h, the temperature was raised at 20 K/min, and the heating interval was 100 K.

#### **2 RESULTS AND DISCUSSION**

## **2.1 Synthetic Ferrosilite**

Under the conditions of different pressures and the same temperature, the run products are completely different. At 1 GPa and 1 273 K (Fig. 2c, YJ-170), the reaction products is mainly composed of fayalite and quartz; when the pressure is increased to 2–3 GPa, the reaction product is ferrosilite (Figs. 2b and 2e, YJ-56 and YJ-55). Obviously, as the pressure in‐ creases, ferrosilite can be better sintered under the same conditions. The molar volumes of fayalite, quartz, and ferrosilite are 46.41, 22.68, and 33.44 cm<sup>3</sup>/mol, respectively. The sum of molar volumes of unit fayalite and quartz is greater than that of two unit ferrosilite (Table 3, modified from Lindsley et al., 1964). At higher pressures the reaction (1) will proceed towards the ferrosilite with a lower molar volume.

Under the same pressure and different temperatures, the run products are different. For example, at 2 GPa the run prod‐ ucts are ferrosilite at 1 273 K (Fig. 2e, YJ-55) and Fa + Qtz at 1 373 K (Fig. 2f, YJ-58), respectively. Therefore, 2 GPa and 1 323 K may be the boundary conditions of the phase transi‐ tion between Fs and Fa + Qtz. At constant pressure, higher temperature will increase the energy of the system, and enhance the vibration of the particles. Consequently, the reaction (1) will proceed towards fayalite and quartz with a larger molar volume (Newnham, 1975).



T(K)	$\kappa$ (W · m <sup>-1</sup> · K <sup>-1</sup> )	$D \text{ (mm}^2 \text{·s}^{-1})$	$c_p$ (kJ·kg <sup>-1</sup> ·K <sup>-1</sup> )
292.24	4.275(214)	1.383(69)	0.785(55)
374.05	4.001(200)	1.120(56)	0.906(63)
476.72	3.637(184)	0.975(48)	0.947(66)
573.86	3.224(114)	0.870(44)	0.941(66)
674.35	3.098(161)	0.788(39)	0.997(70)
774.02	2.946 (147)	0.728(36)	1.027(72)
875.41	2.965(148)	0.704(35)	1.069(75)

**Table 3** Molar volume of mineral aggregate (modified after Lindsley et al., 1964)



 $SiO<sub>2</sub>$  with nanometer grain size is more efficient than micrometer grain size in synthesizing ferrosilite due to the low re‐ action dynamics of SiO<sub>2</sub>. To facilitate the reaction Fe<sub>2</sub>O<sub>3</sub> + Fe +  $3SiO<sub>2</sub> = 3FeSiO<sub>3</sub>$ , Xu et al. (2020) added 1 wt.% water to the starting materials, which inevitably cause the contamination of sample by hydrogen. By reducing the grain size of  $SiO<sub>2</sub>$ , ferrosilite can be easily synthesized with high quality. Among the three types of capsules, graphite, Fe, and Mo, the graphite cap‐ sule can provide a reducing environment for synthesis experiments and does not react with the starting materials. In summary, the ideal conditions for synthesizing ferrosilite are 3 GPa, 1 273 K, nano-scale SiO, as the starting material, and graphite as the sample capsule.

High porosity will significantly reduce the *κ* and *D* of polycrystalline samples (Xiong et al., 2020). In addition, high porosity of the sample will cause the difficulty in machining, larger deformation of the sample under high temperature and high pressure, larger temperature gradient, lower signal quality,



**Figure 2.** Back-scattered images of the recovered samples. The images of (a) synthetic ferrosilite; (b) and (e) synthetic samples under different pressures; (c), (e) and (f) synthetic samples at different temperatures; (b) and (d) the images of synthetic samples using micro-scale SiO, and nano-scale as the starting materials, respectively; (g), (h), (i) and (j) the images of synthetic samples using different capsules.



Figure 3. Ferrosilite sample with a diameter of 5 mm and a thickness of 1 mm. be the dominant mechanism.

#### **2.2 Thermal Conductivity and Thermal Diffusivity**

Figure 5 shows the signal of thermal disturbance generated by one single pulse heating (100 ms, 12 V). It clearly demonstrates the temperature change caused by the heat conduc‐ tion of the sample. Using equations (2) and (3) to fit the ther‐ mal disturbance signal, we can obtain the thermal conductivity and thermal diffusion coefficient of ferrosilite at different temperatures. The fitting errors of parameters *A* and *B* are 1.26% and 1.92%, respectively. The relationship between the *κ* and *D* and the temperature can be expressed by the empirical formulas

$$
\kappa = a_0 + a_1 T^{-1} + a_2 T^{-2} \tag{4}
$$

$$
D = b_0 + b_1 T^{-1} + b_2 T^{-2}
$$
 (5)

where  $a_0$ ,  $a_1$ ,  $a_2$ ,  $b_0$ ,  $b_1$  and  $b_2$  are the constants which we could obtained from non-linear least square fitting. The fitting results yield  $a_0 = 1.786$ ,  $a_1 = 1.048 \times 10^3$ ,  $a_2 = -9.269 \times 10^4$ ,  $b_0 = 0.424$ ,  $b_1 = 0.223 \times 10^3$  and  $b_2 = 1.640 \times 10^4$  (Table 4). Both the thermal conductivity and the thermal diffusivity decrease with increasing temperature, indicating the phonon conduction may



**Figure 4.** Schematic diagram of the cell assembly for the thermal properties measurements.



**Figure 5.** The records of oscilloscope of ferrosilite. (a) CH1 and CH2 are the signals of the inpulse heater and the thermocouple; (b) the amplified signals.

$\kappa(T) = a_0 + a_1 T^{-1} + a_2 T^{-2}$								
$a_0$ (W·m <sup>-1</sup> ·K <sup>-1</sup> )	$a_{1}$ (W · m <sup>-1</sup> )	$a, (W \cdot m^{-1} \cdot K)$	$R^2$					
1.786	$1.048 \times 10^{3}$	$-9.269 \times 10^{4}$	0.980					
$D(T) = b_0 + b_1 T^{-1} + b_2 T^{-2}$								
$b_0$ (mm <sup>2</sup> ·s <sup>-1</sup> )	$b_1$ (mm <sup>2</sup> ·s <sup>-1</sup> ·K)	$b,$ (mm <sup>2</sup> ·s <sup>-1</sup> ·K <sup>2</sup> )	$R^2$					
0.424	$0.223 \times 10^3$	$1.640 \times 10^{4}$	0.998					

**Table 4** Coefficient fitting for thermal conductivity and thermal diffusivity as a function of temperature

Since there is no literature data of *κ* and *D* of ferrosilite, we compare the result of this study with natural enstatite En99 (Hof‐ meister, 2012), En98 (Hunt et al., 2011) and synthesized olivine, Fo and Fa (Zhang Y Y et al., 2019) (Figs. 6a and 6b). Thermal conductivity and thermal diffusivity of these samples all de‐ crease with increasing temperature at constant pressure. Ther‐ mal conductivity of En99 reported by Hofmeister (2012) is higher than that of ferrosilite at 1 GPa and temperatures lower than 500 K. Thermal diffusivity of En99 at 1 atm is in agreement with that of ferrosilite at 1 GPa. However, the thermal diffusivities of En98 at 4 and 6 GPa (Hunt et al., 2011) are much higher than that of En99 at 1 atm (Fig. 6b), indicating a strong effect of pressure on the thermal diffusivity. At temperatures higher than 500 K, *κ* and *D* of ferrosilite and En99 become similar. The rela‐ tively large difference between this study and Hofmeister (2012) at low temperatures are similar to that of olivine. In the previous study of Zhang Y Y et al. (2019), the *κ* and *D* of forster‐ ite are much larger than those of olivine with Fo number smaller than 90 at low temperatures. The pressure dependence of *κ* and *D* of forsterite and fayalite [(∂*κ)*/∂*P*, (∂*D)*/∂*P*] are 0.289, 0.122 and 0.064, 0.027, respectively (Zhang Y Y et al., 2019). The pressure has a greater effect on forsterite than fayalite.

The  $\kappa$  and  $D$  of the silicate minerals increase with increasing pressure (Hofmeister, 2007; Gibert et al., 2003), and de‐ crease with increasing temperature and iron content (Zhang Y Y et al., 2019; Hofmeister and Pertermann, 2008; Schatz and Simmons, 1972). Combining the effects of pressure and iron content on the  $\kappa$  and  $D$ , the results of this study are reasonable.

## **2.3 Heat Capacity**

Thermal conductivity *κ* and thermal diffusivity *D* of ferro‐ silite are determined simultaneously using present method, the heat capacity  $(c_n)$  can be calculated using the formula as follow (Clauser, 2011)

$$
c_p = \frac{\kappa}{D\rho} \tag{6}
$$

where  $\rho$  is the density, with that obtained by calorimetric measurements. Figure 6c shows the temperature dependence of the heat capacity  $(c_n)$  of ferrosilite at 1 GPa and the comparison between present results with previous ones (Saxena and Shen, 1992). The calculated heat capacity of ferrosilite at 1 GPa in‐ creases with temperature with about 10% per 100 K  $(<500 K)$ and 4% per 100 K (>500 K). The pressure dependence of the specific heat capacity of ferrosilite is about 10% per GPa, which is larger than that of omphacite, diopside and jadeite (Wang et al., 2014). However, according to error transfer, the



**Figure 6.** The temperature dependence on the thermal physical properties of ferrosilite and the comparison with previous studies. (a) The thermal con‐ ductivity of ferrosilite at 293–873 K. Reference data are from En99 (Hof‐ meister, 2012) and Fa, Fo (Zhang Y Y et al., 2019). (b) The thermal diffusivity of ferrosilite at 293–873 K. Reference data are from En99, En98 (Hof‐ meister, 2012; Hunt et al., 2011) and Fa, Fo (Zhang Y Y et al., 2019). (c) Calculated heat capacity of ferrosilite at 1 GPa and measured one at 1 atm by calorimetric method (Saxena and Shen, 1992).

errors attached to the calculated specific heat capacity data points are about 7% for ferrosilite. Therefore, the difference of the heat capacity between this study and calorimetric measurements is acceptable.

#### **2.4 Planetary Applications**

The thermal history of asteroids plays an important role in understanding terrestrial planet formation. Zhang Y Y et al. (2019) survey the effect of Fe on the cooling history of aster‐ oids. According to their results the cooling rate of asteroids de‐ creases with increasing iron content. In their calibration, they used olivine to simulate the bulk composition. In this study, we used pyroxene instead of olivine to simulate the bulk composition. We used a simple heat conduction model of asteroids (Zhang Y Y et al., 2019). This model completely ignores other complex factors related to the thermal evolution of asteroids, such as the role of internal heat sources. At the same time, the model does not consider the influence of the shape of the asteroid. Asteroids consist of only a single phase, and only transfer heat through thermal conduction. The radius of the asteroid is 500 km. Initially, the temperature is homogeneously distributed inside the asteroid, with a temperature of 1 500 K. Based on the above assumptions, we use the software COMSOLTM to simulate the heat transfer process. We performed the simulation using constant thermal conductivity at room temperature, which maximizes the difference between different minerals.

Figure 7 show the temperature distribution inside the asteroids at 1 Gy and time-dependent temperature of the center point of the asteroids with constant thermal conductivity. The enstatite asteroid cools very quickly, and it has basically cooled within 1 Gy. The cooling of ferrosilite is slower than enstatite,



Figure 7. Temperature distribution inside the asteroids with different composition model. (a) Temperature distribution inside asteroids at 1 Gy. (b) Time-dependent temperature at the center of the asteroids with constant thermal conductivity. Reference data are from En99 at 1 atm (Hofmeister, 2012) and Fa, Fo at 1 GPa (Zhang Y Y et al., 2019).

and the temperature only drops by 66% at 1 Gy. The temperature between the internal and surface of the enstatite asteroid is less than 100 K, while the ferrosilite asteroid is greater than 500 K. This is basically consistent with the simulation results of ol‐ ivine endmember minerals. Interestingly, the evolution results of the two asteroids composed of enstatite and forsterite, respec‐ tively, are basically the same, while the two asteroids composed of ferrosilite and fayalite, respectively, have a large gap. The temperature difference between the ferrosilite asteroid and the fayalite asteroid increases with time, and the difference is about 20% at 1 Gy. In order to quantitatively constrain the effect of iron content on the thermal evolution history of asteroids consisting of pyroxene, more experiments with variable iron contents should be conducted in the future. If we consider the asteroids, e.g., the Moon, the composition model consisting of both olivine and pyroxene should be used, in which case the calculat‐ ed values of temperature distribution and time-dependent temperature should be intermediate between the values from olivine model and pyroxene model.

## **3 CONCLUSION**

Based on 12 synthesis experiments, we found that high temperature or low pressure is not benefit of the formation of ferrosilite. We used the sample synthesized for the *in*-*situ* mea‐ surements of the thermal conductivity and the thermal diffusivity. The thermal conductivity and the thermal diffusivity of fer‐ rosilite are consistent with previous studies. The iron content of the bulk composition of asteroids strongly influences their cooling history and the temperature distribution. For Fe-rich as‐ teroids, the mantle will be warmer. As a contrast, the heat will easily lose in the Fe-poor asteroid (e.g., Mercury).

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